

# Adaptive Sliding Mode Control for Asteroid Hovering with Solar Sailing

Application to 433 Eros

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## ➤ Introduction

## ➤ Dynamics

- Dynamics in Cylindrical Form

## ➤ Control

- Design of Sliding Surface
- Design of Controller
- Design of Adaptive Estimation

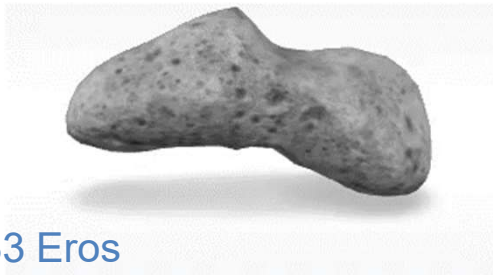
## ➤ Results

- Simulation of Hovering Orbit
- Robustness
- Effect of Hovering Radii and Height
- Effect of Sunlight Direction

## ➤ Conclusions

## Why do we explore asteroids?

- origin of the solar system, planets and life
- space resources extraction
- planetary defence
- ...



433 Eros



25143 Itokawa

Image credit: NASA asteroids, comets & meteors

## NEAR-Shoemaker



## Hayabusa



## Will solar sailing make a difference? **YES! It's propellant-free!**

- It takes great advantage in single/multiple asteroids rendezvous missions.

## Sail operation in proximity of asteroids needs to be studied:

- to maximise scientific return of rendezvous missions
- to support energy-consuming mapping operations, such as hovering
- ...

## Difficulties:

- **Underactuated**: control force constrained in both magnitude and direction.
- **Non-affine**: attitude angles as input is not linear.
- **Complex gravity**: non-spherical perturbation hard to model prior to a mission.

## Our idea:

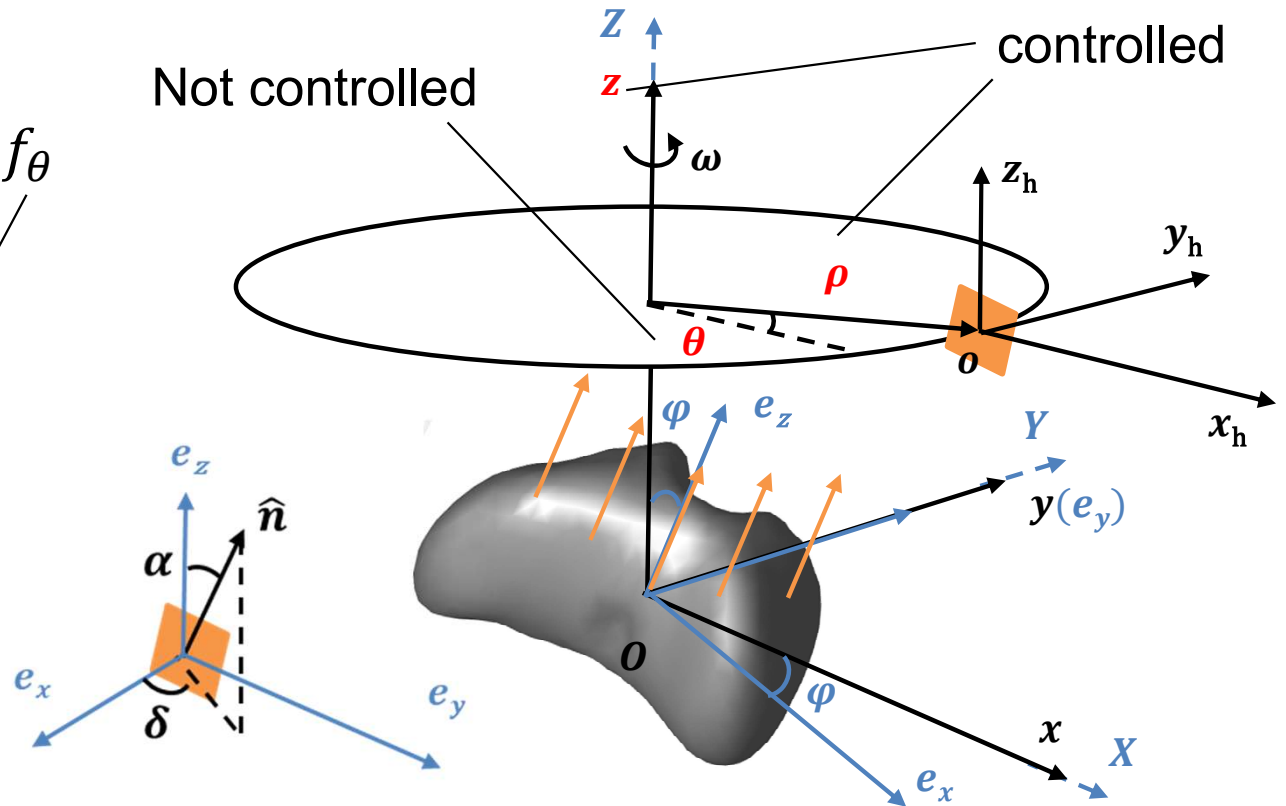
- ✓ To reduce one DOF to be controlled. → **Underactuated**
- ✓ To control the derivatives of attitude angles. → **Non-affine**
- ✓ To use robust control with adaptive estimation. → **Complex gravity**

## Dynamics in cylindrical coordinates $\chi = [\rho, \theta, z]^T$ in asteroid-fixed frame

$$\begin{cases} \ddot{\rho} = \rho(\omega + \dot{\theta})^2 + g_\rho + f_\rho \\ \rho\ddot{\theta} = -2\dot{\rho}(\omega + \dot{\theta}) + g_\theta + f_\theta \\ \ddot{z} = g_z + f_z \end{cases}$$

$$\mathbf{g} = [g_\rho, g_\theta, g_z]^T = \mathbf{C}_a^h \nabla U$$

$$\mathbf{f} = [f_\rho, f_\theta, f_z]^T = \mathbf{C}_I^h \mathbf{C}_E^I \mathbf{a}_{SRP}$$

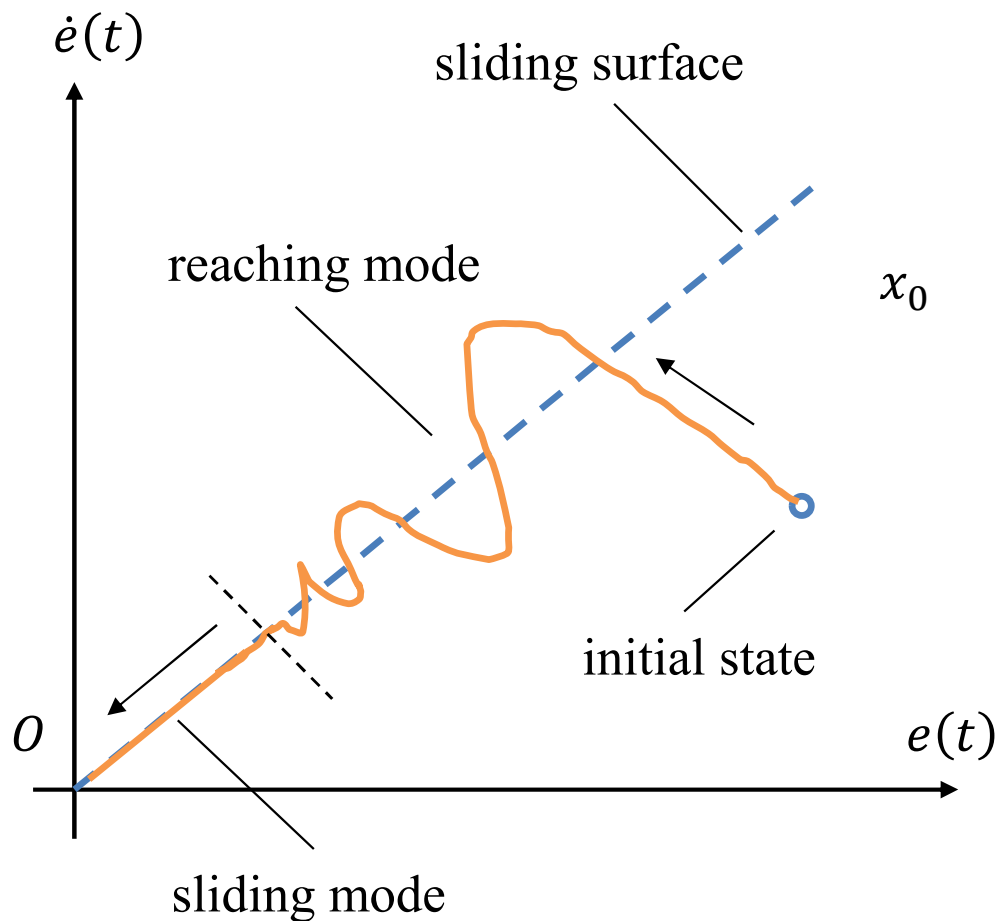


**Objective of control:**

➤ **To track constant  $\rho$  and  $z$ , leaving  $\theta$  aside.**

$$\mathbf{E} \xrightarrow{\mathbf{C}_y(\varphi)} \mathbf{I} \xrightarrow{\mathbf{C}_z(\omega t)} \mathbf{a} \xrightarrow{\mathbf{C}_z(\theta)} \mathbf{h}$$

## Sliding mode control



Tracking error

$$e = \chi - \chi_d = \begin{bmatrix} \rho - \rho_d \\ z - z_d \end{bmatrix}$$

Sliding surface

$$s = \dot{e} + \mathbf{k}e$$

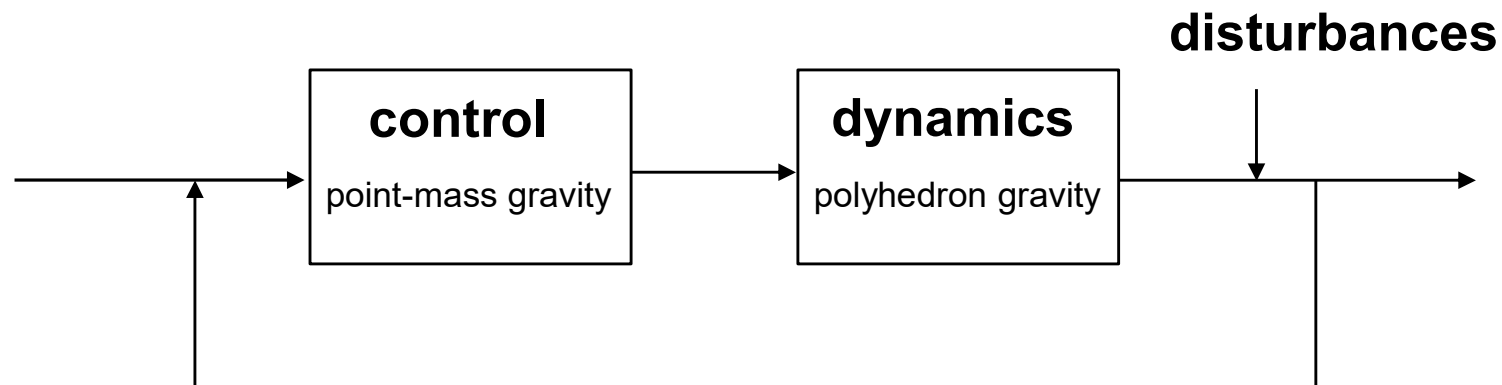
Non-singular terminal sliding surface

$$\sigma = s + k_0 \dot{s}^{\frac{p}{q}}$$

$$\sigma = 0 \rightarrow s = 0 \rightarrow e = 0$$

- $\mathbf{k}, k_0$  – positive numbers
- $p, q$  – positive odd number,  $1 < p/q < 2$

Gravity used is different in control and dynamics!



polyhedron gravity = mass point gravity + disturbances

irregular-shape perturbation / high-order harmonics

## Objective of control:

- **To control  $\dot{\mathbf{u}} = [\dot{\alpha}, \dot{\delta}]^T$  instead of  $\mathbf{u} = [\alpha, \delta]^T$ .**

Further differentiate the dynamics

$$\ddot{\chi} = \mathbf{h}(\chi, \dot{\chi}) + \mathbf{C}_I^o \mathbf{C}_E^I \mathbf{a}_{SRP} + \mathbf{C}_I^o \mathbf{C}_E^I \mathbf{B}(\mathbf{u}) \dot{\mathbf{u}}$$

$$\ddot{\chi} = \begin{cases} \rho(\omega + \dot{\theta})^2 + g_\rho + f_\rho \\ \underline{g_z} + f_z \end{cases}$$

$$\mathbf{h} = \mathbf{h}_0 + \mathbf{d}$$

known portion

unknown disturbance  
(by irregular gravity)

Assumption:

$$\|\mathbf{d}\| \leq \mathbf{D}$$



Differentiate  $s$  twice and  $\sigma$  once:

$$\ddot{s} = \mathbf{h} + \dot{\mathbf{C}}_I^o \mathbf{C}_E^I \mathbf{a}_{SRP} + \mathbf{C}_I^o \mathbf{C}_E^I \mathbf{B}(\mathbf{u}) \dot{\mathbf{u}} - \ddot{\chi}_d + \mathbf{k}\ddot{\mathbf{e}}$$

$$\dot{\boldsymbol{\sigma}} = k_0 \frac{p}{q} \text{diag}(\dot{s}^{\frac{p}{q}-1}) \left( \frac{q}{kp} \dot{s}^{2-\frac{p}{q}} + \ddot{s} \right)$$

Design reaching law as:

$$\dot{\boldsymbol{\sigma}} = \text{diag}(\dot{s}^{\frac{p}{q}-1}) (-\varepsilon_1 \boldsymbol{\sigma} - \varepsilon_2 \text{sign}(\boldsymbol{\sigma}))$$

|
|  
 rapidity term      robustness term

Control law obtained:

$$\dot{\mathbf{u}} = (\mathbf{C}_I^o \mathbf{C}_E^I \mathbf{B})^{-1} \left[ \ddot{\chi}_d - \mathbf{h} - \dot{\mathbf{C}}_I^o \mathbf{C}_E^I \mathbf{a}_{SRP} - \mathbf{k}\ddot{\mathbf{e}} - \frac{q}{kp} \dot{s}^{2-\frac{p}{q}} - \varepsilon_1 \boldsymbol{\sigma} - \varepsilon_2 \text{sign}(\boldsymbol{\sigma}) \right]$$

## Objective of control:

- **To be robust against gravity disturbances.**

Adaptive estimation on boundary of gravity disturbances:

$$\hat{\mathbf{D}} = \gamma \frac{k_0 p}{q} \text{diag}(\dot{\mathbf{s}}^{q-1}) |\boldsymbol{\sigma}|$$

$$\dot{\mathbf{u}} = (\mathbf{C}_I^o \mathbf{C}_E^I \mathbf{B})^{-1} \left[ \ddot{\boldsymbol{\chi}}_d - \underline{\mathbf{h}} - \dot{\mathbf{C}}_I^o \mathbf{C}_E^I \mathbf{a}_{SRP} - k \ddot{\mathbf{e}} - \frac{q}{kp} \dot{\mathbf{s}}^{2-\frac{p}{q}} - \varepsilon_1 \boldsymbol{\sigma} - \varepsilon_2 \text{sign}(\boldsymbol{\sigma}) \right]$$

$$\Rightarrow \dot{\mathbf{u}} = (\mathbf{C}_I^o \mathbf{C}_E^I \mathbf{B})^{-1} \left[ \ddot{\boldsymbol{\chi}}_d - \underline{\mathbf{h}_0} - \hat{\mathbf{D}} - \dot{\mathbf{C}}_I^o \mathbf{C}_E^I \mathbf{a}_{SRP} - k \ddot{\mathbf{e}} - \frac{q}{kp} \dot{\mathbf{s}}^{2-\frac{p}{q}} - \varepsilon_1 \boldsymbol{\sigma} - \varepsilon_2 \text{sign}(\boldsymbol{\sigma}) \right]$$

# Results

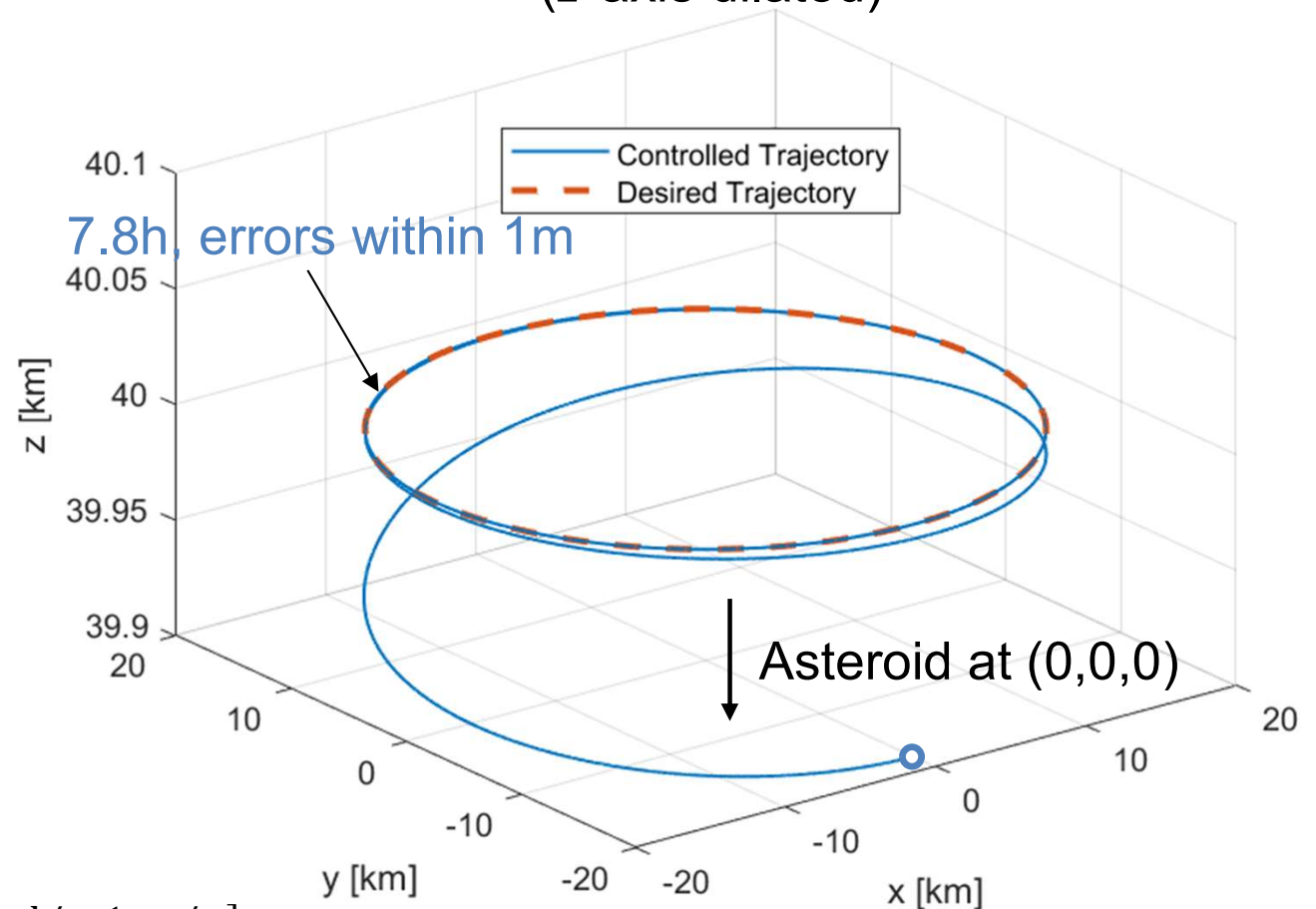
Constants	Value
Eros Gravitational Constant $\mu$	$4.4602 \times 10^4 \text{ km}^3/\text{s}^2$
Eros dimension	$34.4 \times 11.2 \times 11.2 \text{ km}$
Eros Spin Rate $\omega$	$3.3117 \times 10^{-4} \text{ rad/s}$
Eros Heliocentric Distance $R$	$1.6917 \times 10^6 \text{ km}$
Solar Incidence Angle $\varphi$	0 deg
Sail Lightness Number $\beta$	0.2

Initial conditions:

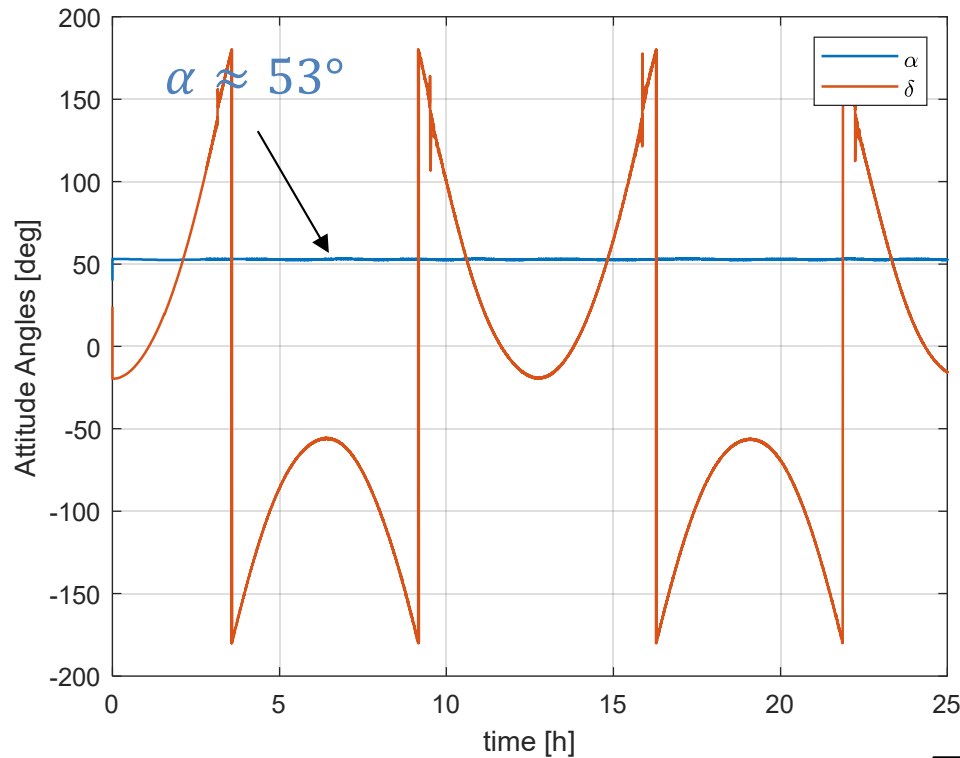
$$[\rho, \theta, z]^T = [18.1 \text{ km}, -\pi/2, 39.9 \text{ km}]$$

$$[\dot{\rho}, \dot{\theta}, \dot{z}]^T = [-1 \text{ m/s}, -3.3117 \times 10^{-4} \text{ rad/s}, 1 \text{ m/s}]$$

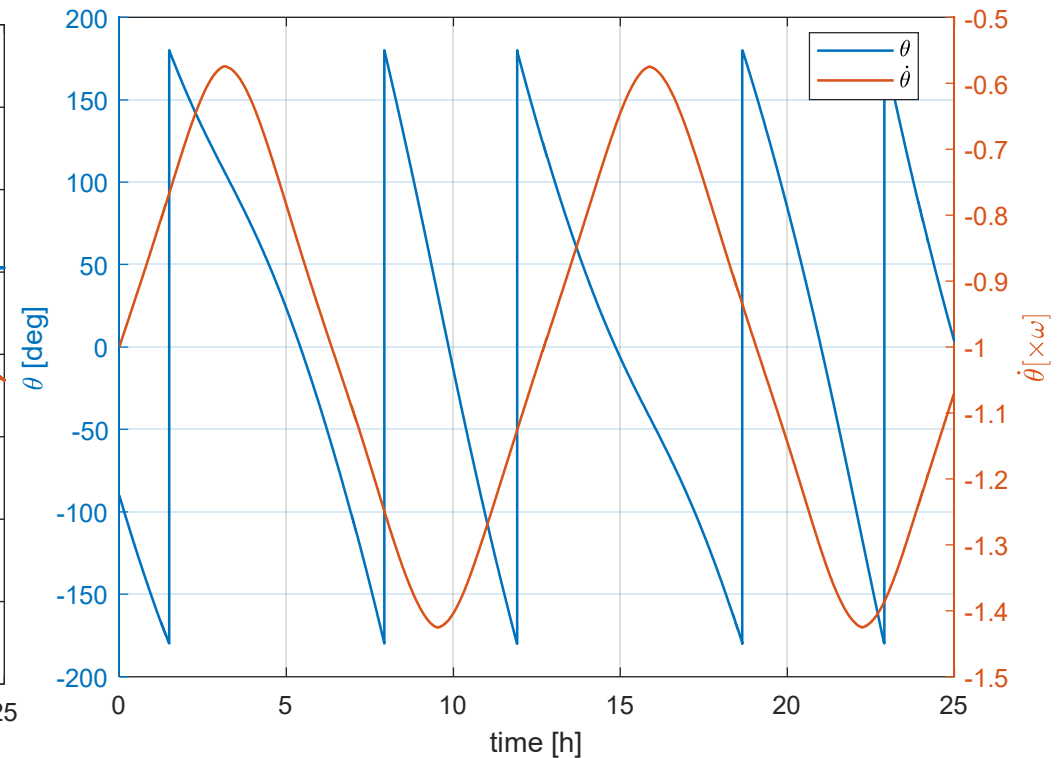
Controlled trajectory to  $\rho_d = 18 \text{ km}$ ,  $z_d = 40 \text{ km}$   
(z-axis dilated)



Control of attitude angles  
 $\alpha$  (red),  $\delta$  (blue)



Uncontrolled polar angle  $\theta$  (blue)  
and polar angular velocity  $\dot{\theta}$  (red)

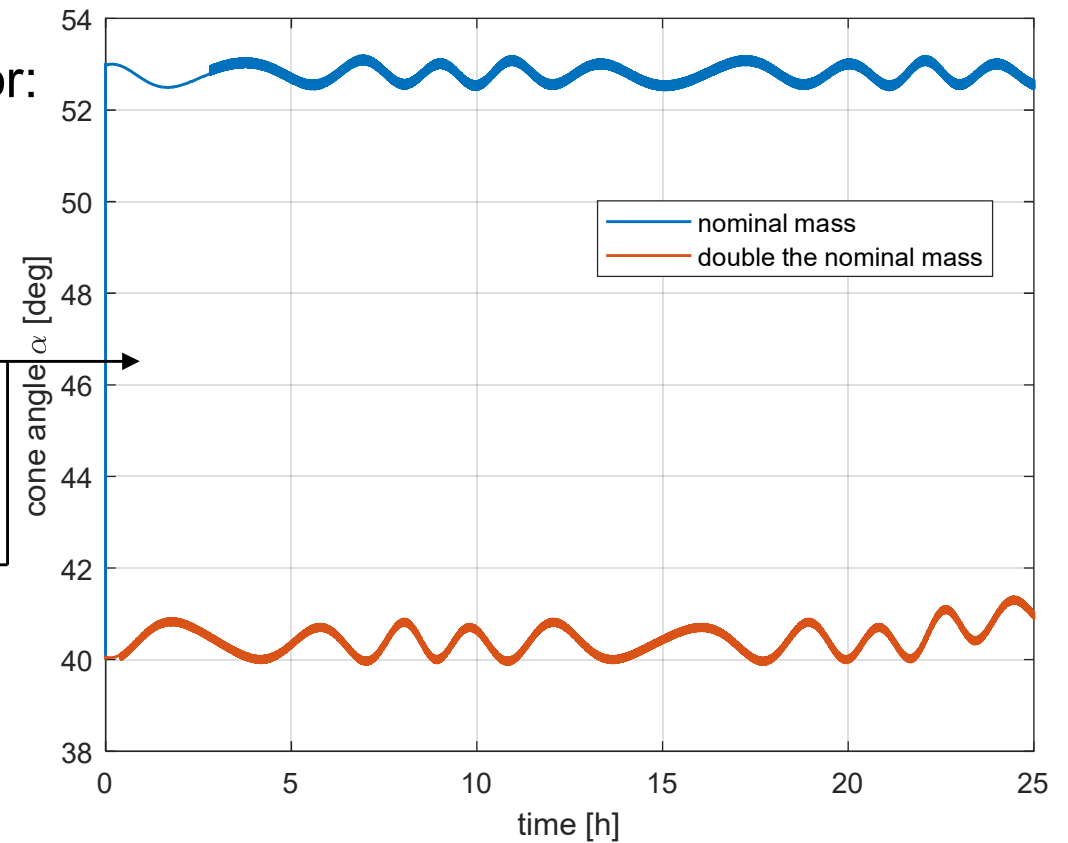
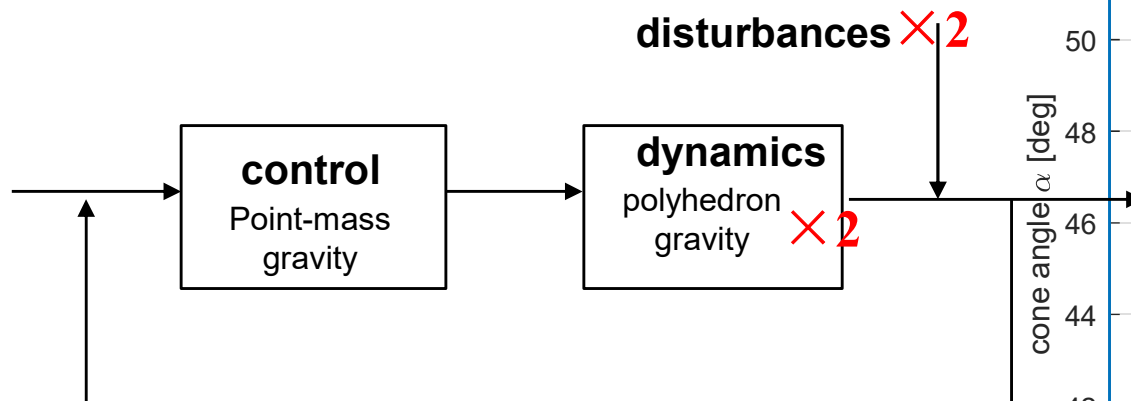


$$\rho \ddot{\theta} = \underbrace{-2\dot{\rho}(\omega + \dot{\theta})}_0 + g_\theta + \underbrace{f_\theta}_{\text{dominant term}}$$

## Robustness

Response to cone angle with nominal asteroid mass (blue) and double nominal mass (red)

Knowledge of asteroid mass with large error:



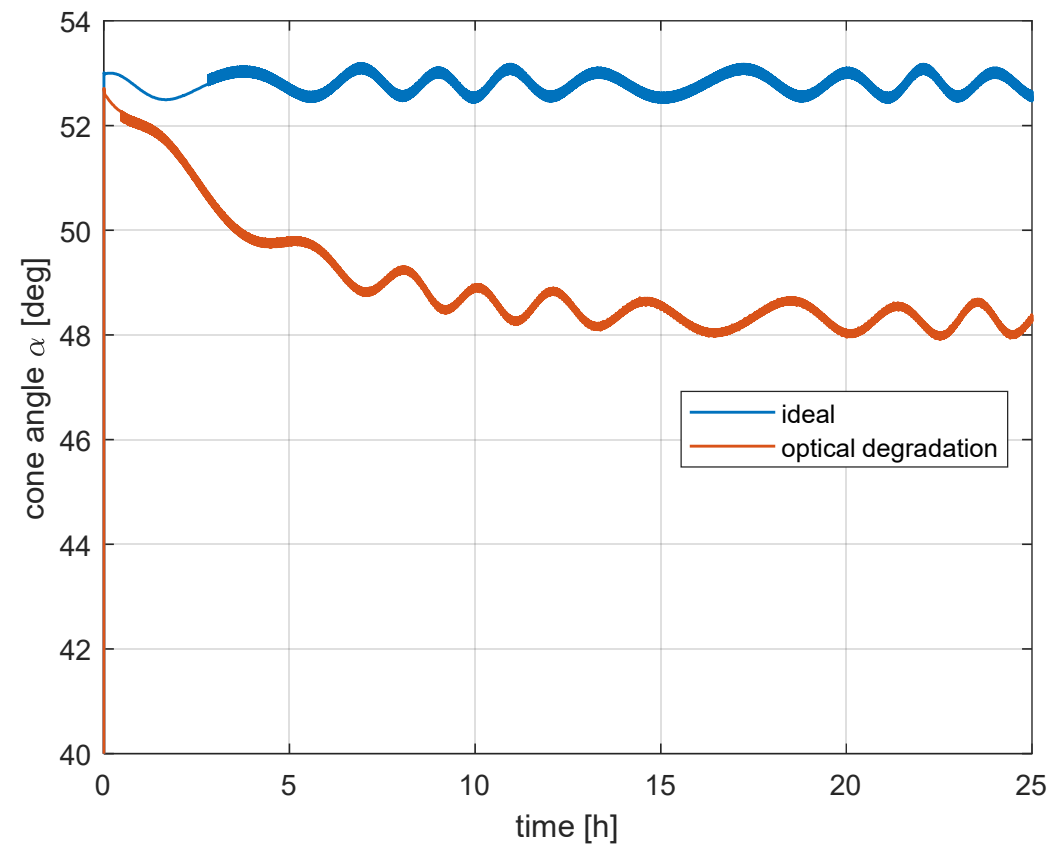
## Robustness

Optical degradation:

$$\beta(t) = 0.05e^{-t/13500} + 0.15$$

$$\beta: 0.2 \xrightarrow{15h} 0.15$$

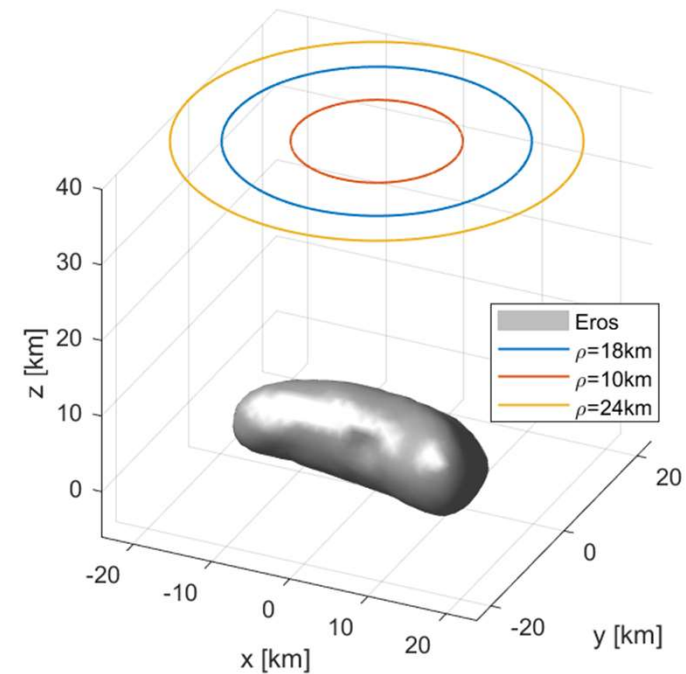
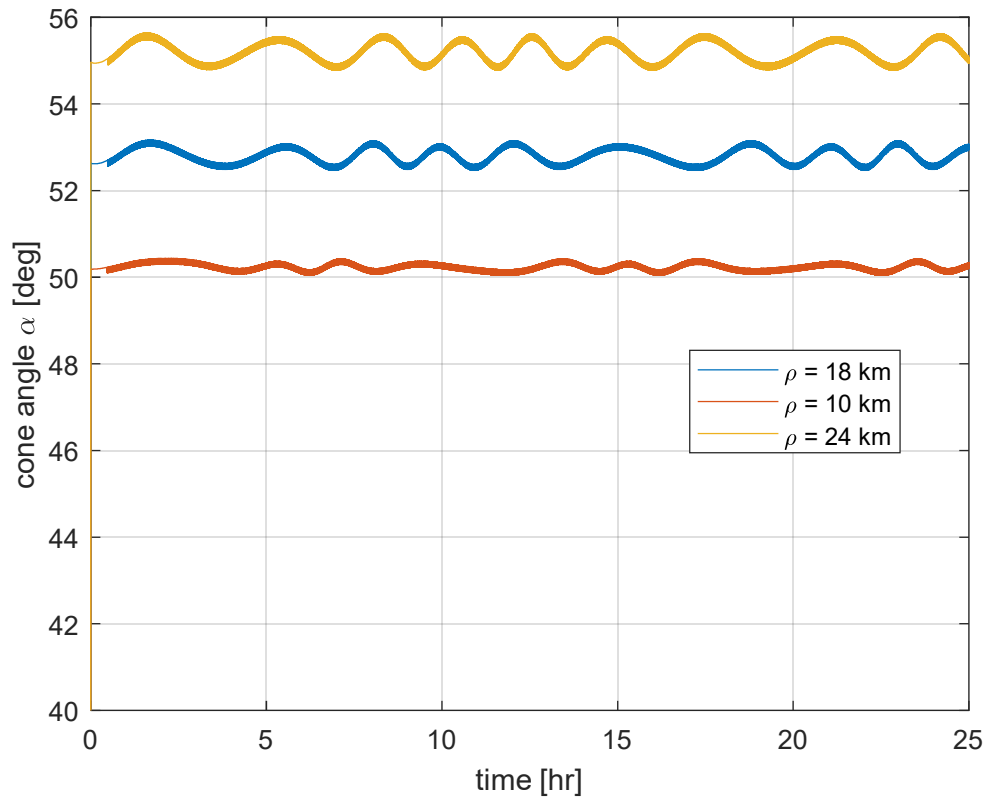
Response to cone angle with ideal sail (blue) and degraded sail (red)



## Effect of Hovering Radii

Response to cone angle with different radii

$z = 40$  km,  $\rho = 18$  km (blue),  $\rho = 10$  km (red),  $\rho = 24$  km (yellow)

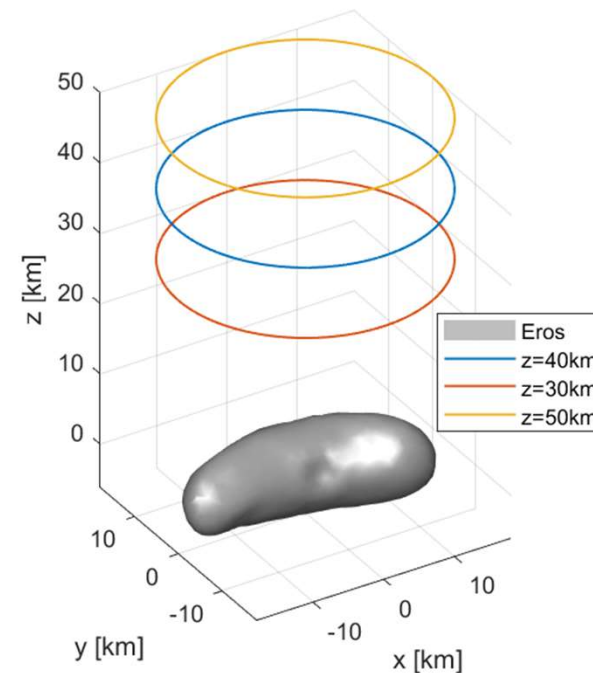
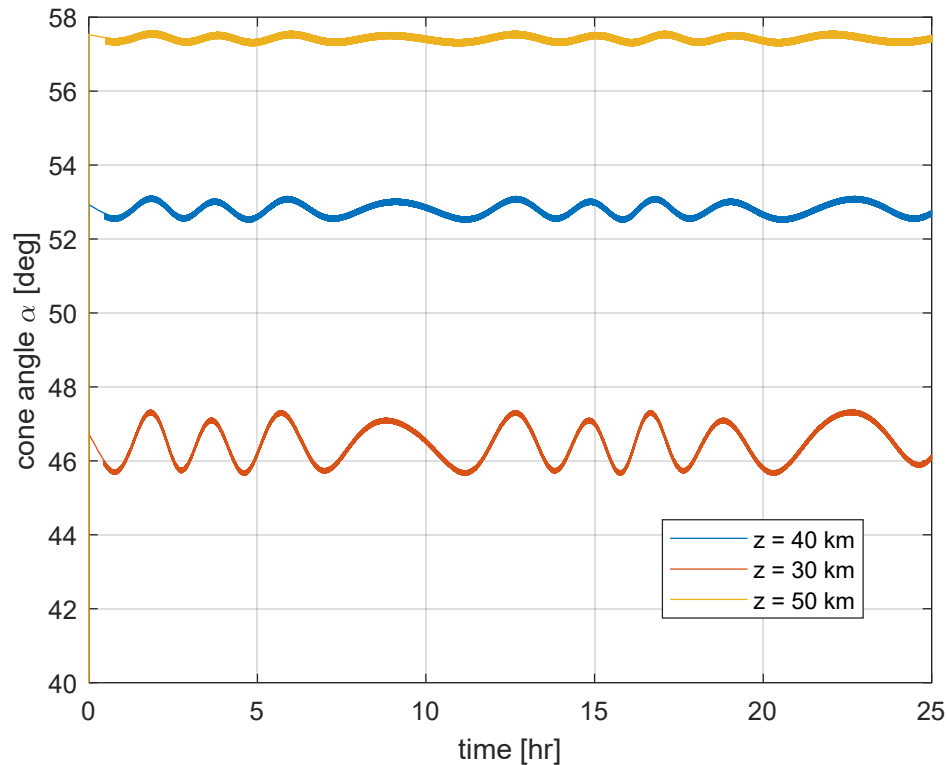


$\rho \nearrow$ ,  $\alpha \nearrow$ , oscillation of  $\alpha \nearrow$

## Effect of Hovering Height

Response to cone angle with different height

$\rho = 18$  km,  $z = 40$  km (blue),  $z = 30$  km (red),  $z = 50$  km (yellow)



$z \nearrow$ ,  $\alpha \nearrow$ , oscillation of  $\alpha \searrow$



## Effect of Sunlight Direction

Sunlight incidence angle is a variable, affected by:

- Eros orbital inclination
- Eros obliquity to the ecliptic

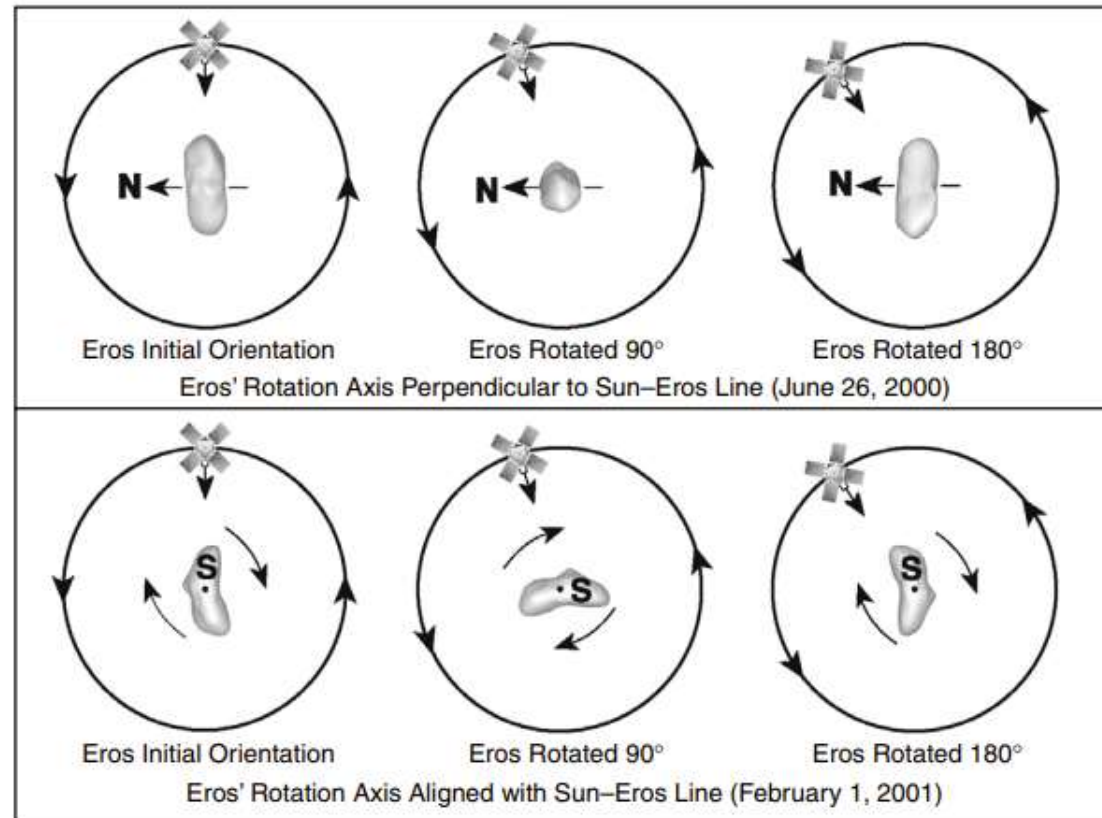
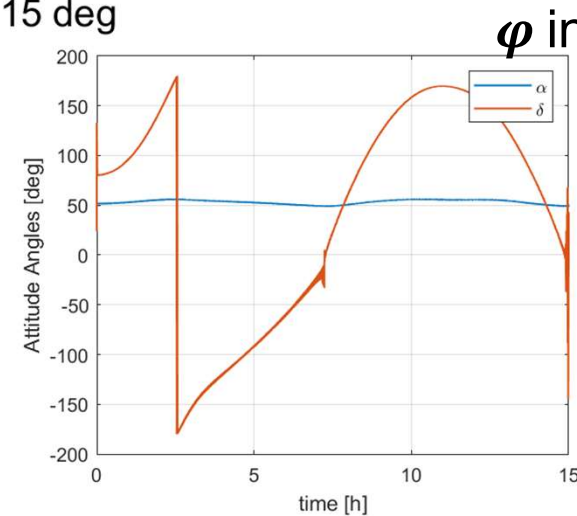
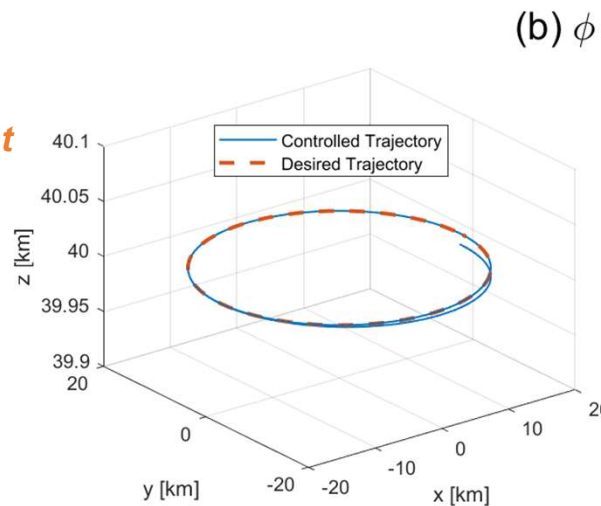
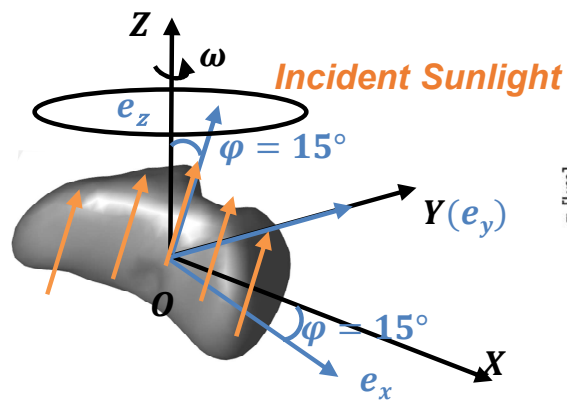
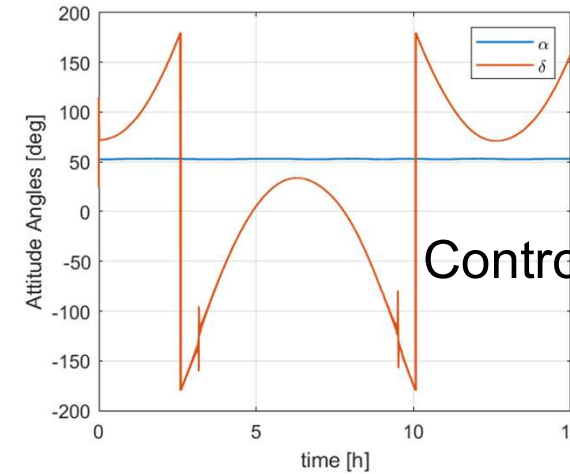
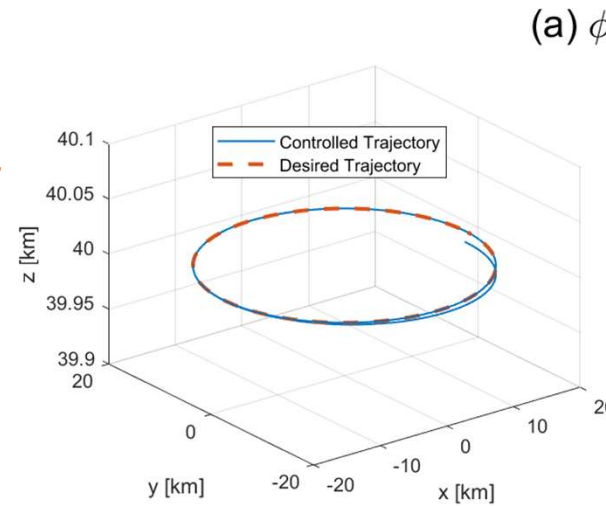
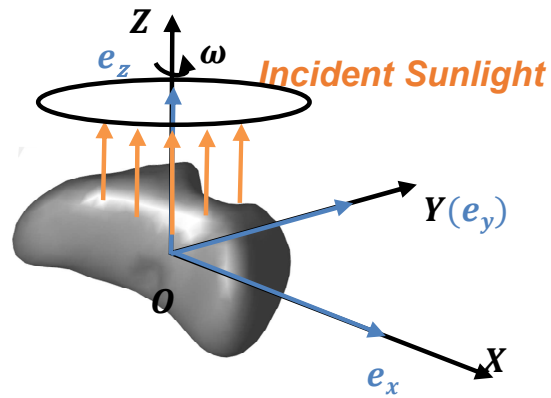
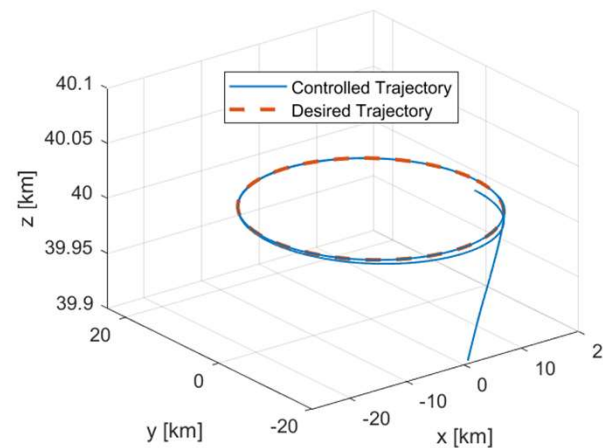
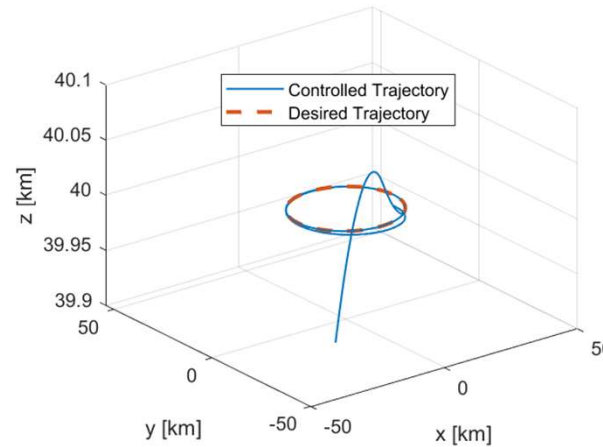
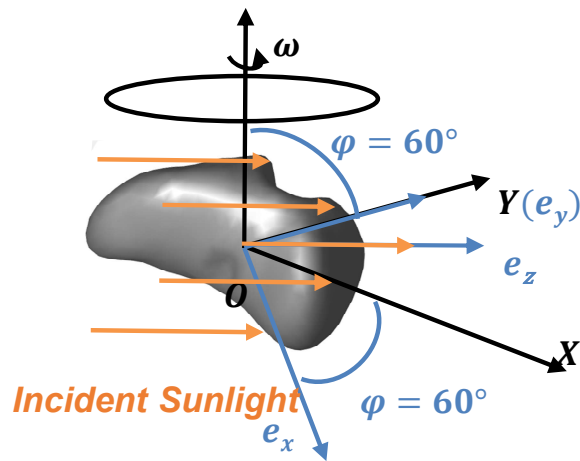
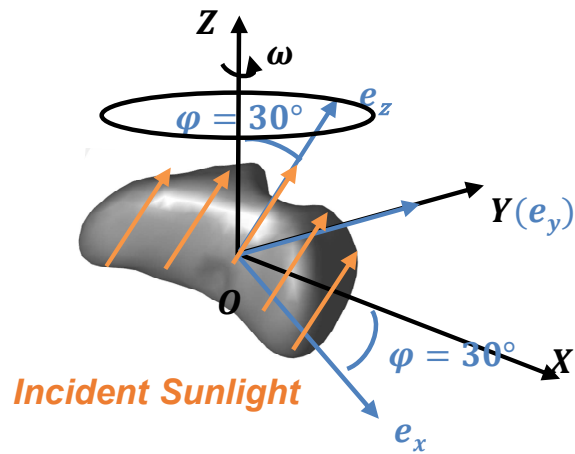


Image credit: Space Exploration of Asteroids: The 2001 Prospective (R. Farquhar)

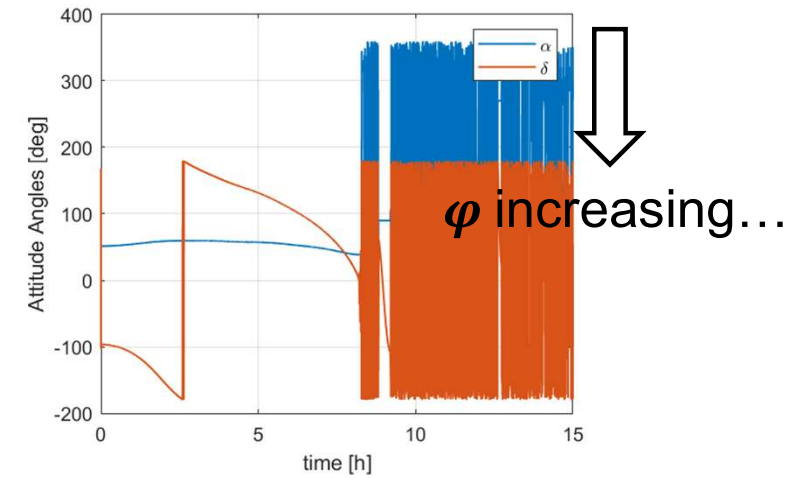
## Effect of Sunlight Direction



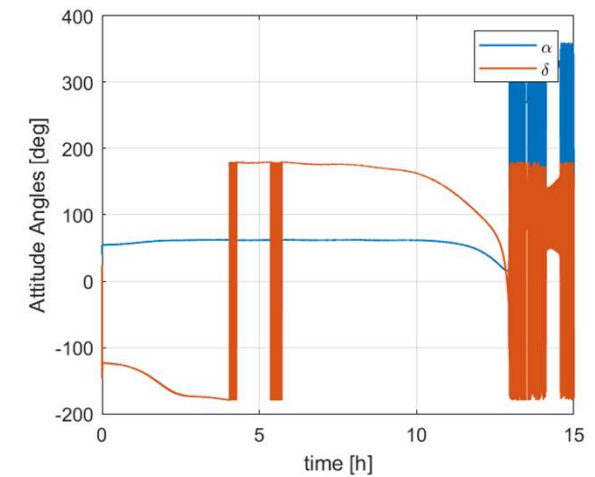
## Effect of Sunlight Direction



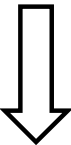
(c)  $\phi = 30$  deg Orbit maintenance failure



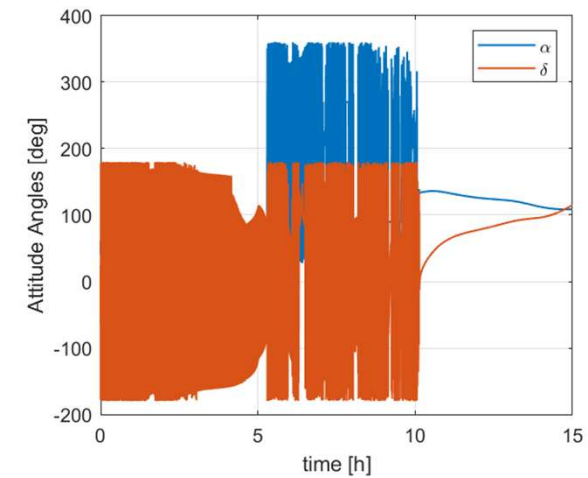
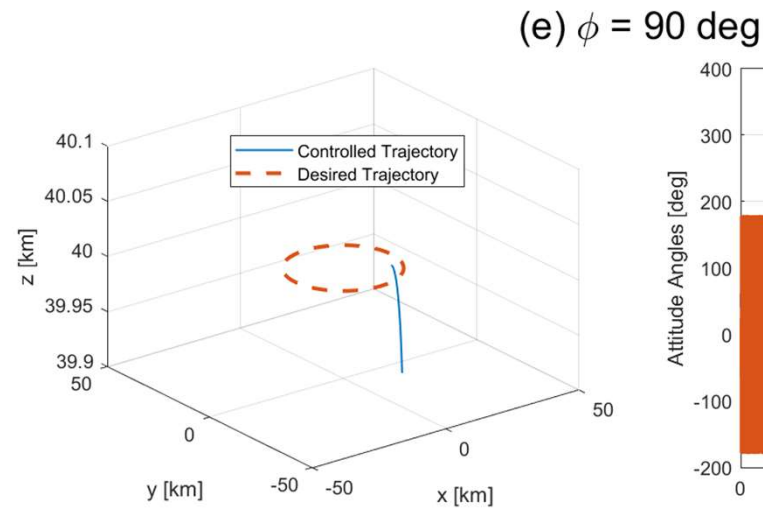
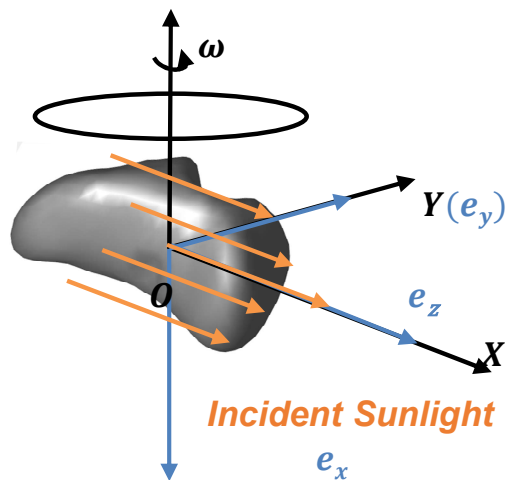
(d)  $\phi = 60$  deg



## Effect of Sunlight Direction

$\phi$  increasing... 

Injection failure



Control breaks down when required control force is sunward.

## Conclusions

- It finds a path to tackle underactuated and non-affine control of solar sail;
- It is robust to gravity disturbances and unmodelled sail error;
- Small hovering radius and height → small cone angle response;
- Small hovering radius and large hovering height → slight cone angle oscillation;
- Control is only effective when sunlight incidence angle is small.

## Future works

- Quantitative research on feasible hovering radius, height and sunlight incidence angles;
- Direct adaptive estimation on gravity property and sail performance, instead of boundary of gravity disturbances.

# Space and Exploration Technology Group

Thank you for listening!



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of Glasgow

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