

Blended Locally-Optimal Control Laws for Space Debris Removal in LEO using a Solar Sail

Christian Bianchi^{1,2}, Lorenzo Niccolai², Giovanni Mengali², Matteo Ceriotti¹

¹James Watt School of Engineering, University of Glasgow, Glasgow, UK

²Department of Civil and Industrial Engineering, University of Pisa, Pisa, Italy

6th International Symposium on Space Sailing
New York City, USA, 5th-9th June 2023



UNIVERSITÀ DI PISA

- 1 Introduction
 - Mission Scenario
- 2 Dynamical Model
 - Equations of motion
 - Perturbing Accelerations
 - Eclipses
- 3 Transfer Strategy
 - 1st Phase: semimajor axis increase
 - 2nd Phase: blended control law
 - Descent Phase
- 4 Numerical Simulations
- 5 Conclusion

Table of Contents

1 Introduction

- Mission Scenario

2 Dynamical Model

- Equations of motion
- Perturbing Accelerations
- Eclipses

3 Transfer Strategy

- 1st Phase: semimajor axis increase
- 2nd Phase: blended control law
- Descent Phase

4 Numerical Simulations

5 Conclusion

Introduction - Space debris

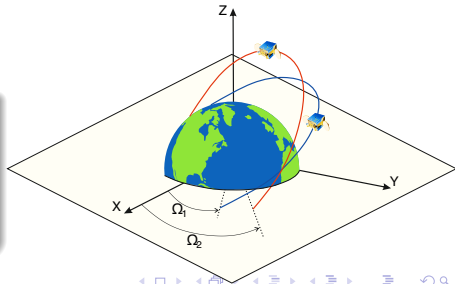
Space Debris Problem

- Increasing **number of objects**
- **Active removal** strategies
- **Constellations**



Space Debris Model

- **Circular** orbits
- **Same inclination**
- **Different RAAN**



Introduction - Mission Scenario

Description

Solar Sail to:

- **reach** debris orbit
- **collect** it
- **bring it down**

Assumption

No phasing (ν is free)

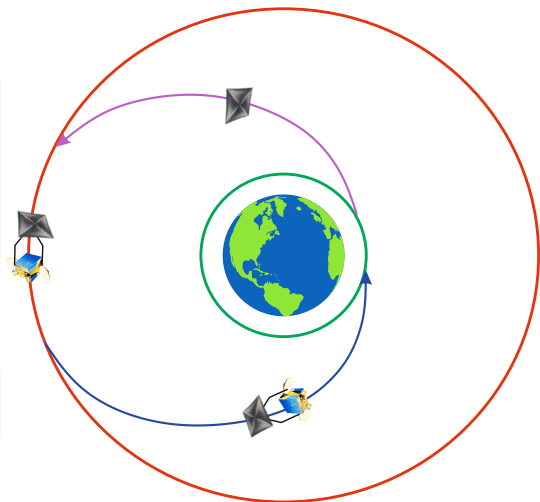


Table of Contents

- 1 Introduction
 - Mission Scenario
- 2 **Dynamical Model**
 - Equations of motion
 - Perturbing Accelerations
 - Eclipses
- 3 Transfer Strategy
 - 1st Phase: semimajor axis increase
 - 2nd Phase: blended control law
 - Descent Phase
- 4 Numerical Simulations
- 5 Conclusion

Equations of Motion

Modified Equinoctial Orbital Elements (MEOEs)

$$p = a(1 - e^2)$$

$$f = e \cos(\Omega + \omega)$$

$$g = e \sin(\Omega + \omega)$$

$$h = \tan(i/2) \cos(\Omega)$$

$$k = \tan(i/2) \sin(\Omega)$$

$$L = \Omega + \omega + \nu$$

$$\mathbf{x} = [p, f, g, h, k, L]^T \quad \text{sail state vector}$$

$$\dot{\mathbf{x}} = \mathbb{A}(\mathbf{x}) \mathbf{a} + \mathbf{b}(\mathbf{x}) \quad \text{equation of motion}$$

$$\mathbf{a} = \mathbf{a}_{SRP} + \mathbf{a}_D + \mathbf{a}_L + \mathbf{a}_{J_2} \quad \text{perturbing acceleration (in RTN frame)}$$

Equations of Motion

$$\mathbb{A}(\mathbf{x}) = \begin{bmatrix} 0 & \frac{2p}{q} \sqrt{\frac{p}{\mu}} & 0 \\ \sqrt{\frac{p}{\mu}} \sin L & \sqrt{\frac{p}{\mu}} \frac{1}{q} [(q+1) \cos L + f] & -\sqrt{\frac{p}{\mu}} \frac{g}{q} [h \sin L - k \cos L] \\ -\sqrt{\frac{p}{\mu}} \cos L & \sqrt{\frac{p}{\mu}} \frac{1}{q} [(q+1) \sin L + g] & \sqrt{\frac{p}{\mu}} \frac{f}{q} [h \sin L - k \cos L] \\ 0 & 0 & \sqrt{\frac{p}{\mu}} \frac{s^2 \cos L}{2q} \\ 0 & 0 & \sqrt{\frac{p}{\mu}} \frac{s^2 \sin L}{2q} \\ 0 & 0 & \sqrt{\frac{p}{\mu}} \frac{1}{q} [h \sin L - k \cos L] \end{bmatrix}$$

$$\mathbf{b}(\mathbf{x}) = \left[0 \quad 0 \quad 0 \quad 0 \quad 0 \quad \sqrt{\mu p} \left(\frac{q}{p} \right)^2 \right]^T$$

$$q = 1 + f \cos L + g \sin L$$

$$s^2 = 1 + h^2 + k^2$$

Solar Radiation Pressure Acceleration

Ideal Force Model

- **Flat** plate
- Only **specular reflection**

$$\mathbf{a}_{SRP} = \eta a_c \cos^2 \alpha \hat{\mathbf{n}}$$

η : **shadow factor** ($\eta = 0$ eclipse, $\eta = 1$ sunlight)

a_c : **characteristic acceleration**

$\alpha \in [0, \pi/2]$: **cone angle**

Assumption

The dependence of \mathbf{a}_{SRP} on the distance from the Sun is neglected

Solar Radiation Pressure Acceleration

Sunlight reference frame \mathcal{T}_S :

$\hat{\mathbf{x}}_S$: Sun-sail direction

$$\hat{\mathbf{y}}_S = \hat{\mathbf{z}}_I \times \hat{\mathbf{x}}_S$$

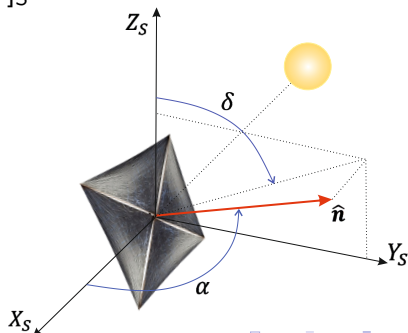
$$[\hat{\mathbf{n}}]_S = \cos \alpha \hat{\mathbf{x}}_S + \sin \alpha \sin \delta \hat{\mathbf{y}}_S + \sin \alpha \cos \delta \hat{\mathbf{z}}_S$$

$$[\hat{\mathbf{n}}]_{RTN} = \mathbb{R}_{S \rightarrow RTN} [\hat{\mathbf{n}}]_S$$

Control Variables

Cone angle $\alpha \in [0, \pi/2]$

Clock angle $\delta \in [0, 2\pi)$



Atmospheric Drag and Lift

Atmospheric Model

- **Hyperthermal** free-molecular flow ($v_{S/C} \gg v_{\text{thermal}}$)
- **NRLMSISE-00** model (MATLAB function *atmosnrlmsise00*)
- **Flat plate**

$$\mathbf{a}_D = -\frac{1}{2} \frac{\rho v^2}{\sigma} C_D \hat{\mathbf{v}}$$

$$\mathbf{a}_L = \frac{1}{2} \frac{\rho v^2}{\sigma} C_L \hat{\mathbf{L}}$$

ρ : **atmospheric density** (*atmosnrlmsise00*)

σ : **sail loading** ($\sigma = m/A$)

v : sail inertial **velocity**

Atmospheric Drag and Lift

Drag and **Lift** coefficients:

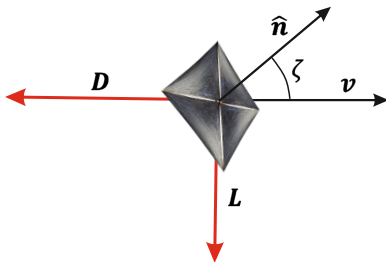
$$C_D = 2 \left[\sigma_T + \sigma_N V_R |\cos \zeta| + (2 - \sigma_N - \sigma_t) \cos^2 \zeta \right] |\cos \zeta|$$

$$C_L = 2 \left[\sigma_N V_R + (2 - \sigma_N - \sigma_T) \cos^2 \zeta \right] |\cos \zeta| \sin \zeta$$

σ_N, σ_T : normal and tangential **accommodation coefficients** (≈ 0.8)

V_R : **ratio** of the average **thermal speed** to the **sail velocity** (≈ 0.05)

$\zeta \in [0, \pi]$: angle between \hat{n} and \hat{v}



Earth's oblateness

$$[\mathbf{a}_{J_2}]_R = -\frac{3\mu J_2 R_{\oplus}^2}{2r^4} \left[1 - \frac{12(h \sin L - k \cos L)^2}{(1 + h^2 + k^2)^2} \right]$$

$$[\mathbf{a}_{J_2}]_T = -\frac{12\mu J_2 R_{\oplus}^2}{r^4} \left[\frac{(h \sin L - k \cos L)(h \cos L + k \sin L)}{(1 + h^2 + k^2)^2} \right]$$

$$[\mathbf{a}_{J_2}]_N = -\frac{6\mu J_2 R_{\oplus}^2}{r^4} \left[\frac{(1 - h^2 - k^2)(h \sin L - k \cos L)}{(1 + h^2 + k^2)^2} \right]$$

$\mu = 3.986 \times 10^5 \text{ km}^3/\text{s}^2$ Earth's **gravitational parameter**

$R_{\oplus} = 6378.14 \text{ km}$ Earth's mean **equatorial radius**

$J_2 = 1.0826 \times 10^{-3}$ Earth's **second harmonic coefficient**

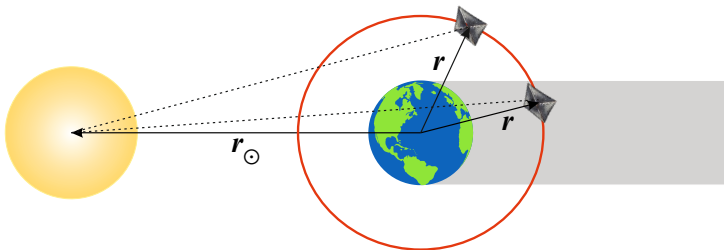
$r =$ **orbital radius** of the sail

Eclipses

Eclipse Model

- **Cylindrical** model

- **Shadow** factor $\eta = \begin{cases} 0 & \text{if } \theta_{\odot} + \theta_{\text{sail}} < \theta \\ 1 & \text{if } \theta_{\odot} + \theta_{\text{sail}} \geq \theta \end{cases} \rightarrow \begin{cases} \text{Eclipse} \\ \text{Sunlight} \end{cases}$



$$\theta = \arccos\left(\frac{\mathbf{r}_{\odot} \cdot \mathbf{r}}{r_{\odot} r}\right), \quad \theta_{\odot} = \arccos\left(\frac{R_{\oplus}}{r_{\odot}}\right), \quad \theta_{\text{sail}} = \arccos\left(\frac{R_{\oplus}}{r}\right)$$

Table of Contents

- 1 Introduction
 - Mission Scenario
- 2 Dynamical Model
 - Equations of motion
 - Perturbing Accelerations
 - Eclipses
- 3 Transfer Strategy**
 - 1st Phase: semimajor axis increase
 - 2nd Phase: blended control law
 - Descent Phase
- 4 Numerical Simulations
- 5 Conclusion

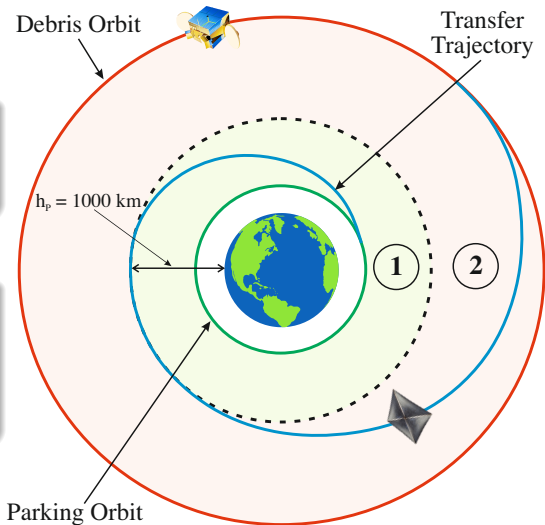
Transfer Strategy

Optimization Technique

Locally-optimal control laws
(orbital perturbations)

Phases

- 1st : SMA increase
- 2nd : debris targeting
- Descent



1st Phase - Description

Objective

Find **optimal sail attitude** $\{\alpha_{opt}, \delta_{opt}\}$ that **minimizes** the **cost index** J at each time (MATLAB *fmincon*)

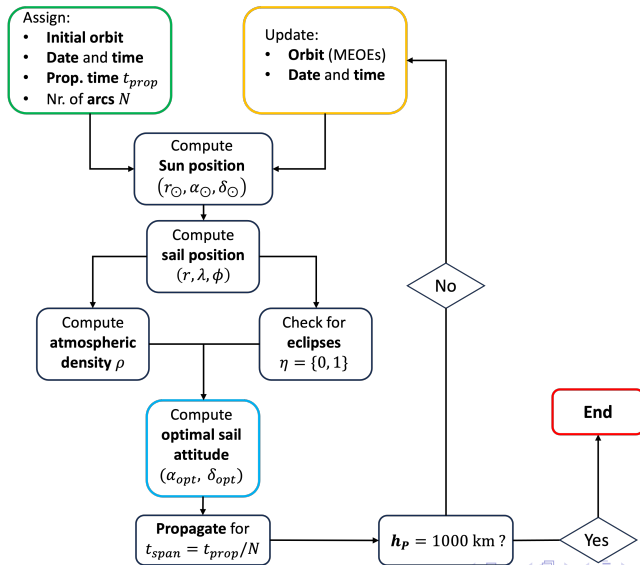
Cost Index

$$J = -\frac{da}{dt} = -2 \sqrt{\frac{a^3}{\mu(1-e^2)}} \left[e \sin \nu a_R + (1 + e \cos \nu) a_T \right]$$

Stopping Criterion

The 1st phase ends when the **perigee altitude** of the osculating orbit reaches 1000 km

1st Phase - Optimization Procedure



2nd Phase - Blended Control Law

Objective

Find **optimal sail attitude** $\{\alpha_{opt}, \delta_{opt}\}$ that **minimizes** the **cost index** J at each time (MATLAB *fmincon*)

Cost Index

$$J = W_a R_a \frac{d(a/a_0)}{dt} + W_e R_e \frac{de}{dt} + W_i R_i \frac{di}{dt}$$

$W_e, W_i \in [0, 1]$: **constant weights**

R_a, R_e, R_i : **variable weights**

$$\frac{de}{dt} = \sqrt{\frac{a(1-e^2)}{\mu}} \left[\sin \nu a_R + \left(\cos \nu + \frac{e + \cos \nu}{1 + e \cos \nu} \right) a_T \right]$$

$$\frac{di}{dt} = \sqrt{\frac{a(1-e^2)}{\mu}} \frac{\cos(\omega + \nu)}{1 + e \cos \nu} a_N$$

2nd Phase - Blended Control Law

Variable Weights

- Adjust the **relative importance** of each orbital element according to the "distance" from the target
- The **sign** indicates if the o.e. time derivative has to be **maximized** ($R < 0$) or **minimized** ($R > 0$)

$$R_a = \frac{a - a_t}{|a_0 - a_t|}, \quad R_e = \frac{e - e_t}{|e_0 - e_t|}, \quad R_i = \frac{i - i_t}{|i_0 - i_t|}$$

where

$\{a_t, e_t, i_t\}$ are the **target semimajor axis, eccentricity and inclination**

$\{a, e, i\}$ are the **sail instantaneous** orbital elements

$\{a_0, e_0, i_0\}$ are the **sail** orbital elements at the **end of the 1st phase**

2nd Phase - RAAN Matching

J₂ Perturbation

- a, e, i undergo short-term oscillations
- Ω has a **secular drift** ($\dot{\Omega} < 0$ for prograde orbits)

Assumption

RAAN behaviour depends only on J_2 (contribution of SRP is negligible)

RAAN of target **debris**: $\Omega_t(t) = \Omega_{t,0} + \dot{\Omega}_t t$

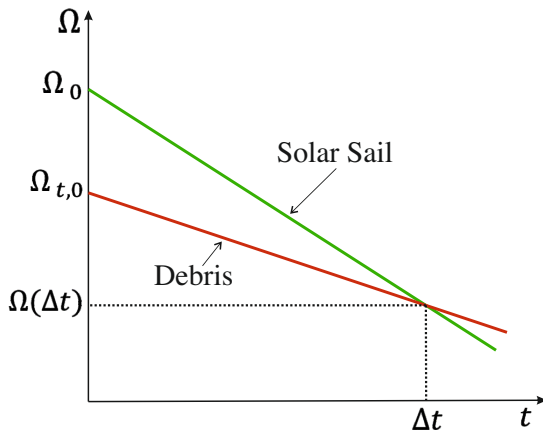
RAAN of **solar sail**: $\Omega(t) = \Omega_0 + \dot{\Omega}_{avg} t$

$$\dot{\Omega}_{avg} = -\frac{3}{2} \frac{J_2 \sqrt{\mu} R_{\oplus}^2}{a_{avg}^{7/2} (1 - e_{avg}^2)^2} \cos i_{avg}$$

$$a_{avg} = (a_0 + a_t)/2 \quad e_{avg} = (e_0 + e_t)/2 \quad i_{avg} = (i_0 + i_t)/2$$

2nd Phase - RAAN Matching

Estimated flight time Δt s.t. $\Omega_t(\Delta t) = \Omega(\Delta t)$



2nd Phase - Genetic Algorithm

Objective

Find the 3 **optimal constant weights** $W_{a,e,i}$ that **minimize** the **objective function** F

GA Objective Function

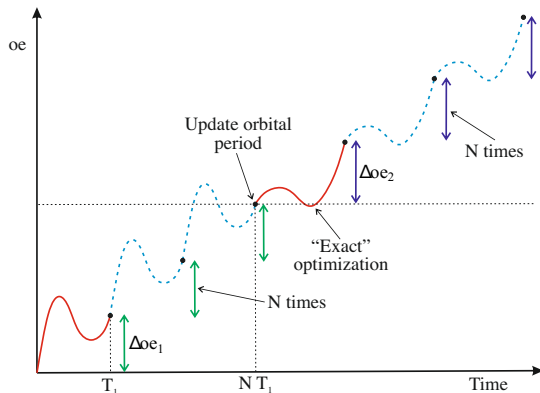
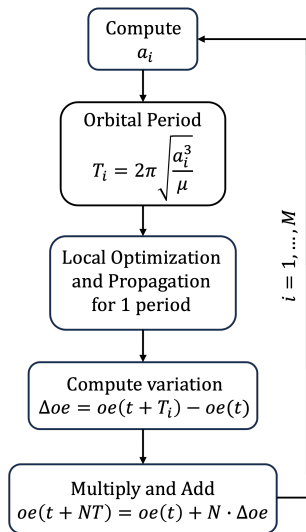
$$F = \left(\frac{a_{\text{fin}} - a_t}{a_0} \right)^2 + (e_{\text{fin}} - e_t)^2 + (i_{\text{fin}} - i_t)^2 + (\Omega_{\text{fin}} - \Omega_t)^2$$

a_{fin} , e_{fin} , i_{fin} , Ω_{fin} : sail orbital elements at the end of propagation time Δt

Approximate Model

An **approximate model** for the transfer is necessary to **reduce the computational time** of the GA

2nd Phase - Approximate Model



Descent Phase - Description

Assumptions

- Flight **time** is **not constrained**
- **RAAN** matching is **not necessary**
- **Sail loading** $\sigma_{\text{desc}} = 2\sigma$ due to **debris mass**
- $\sigma_{\text{desc}} = 2\sigma \implies a_{c,\text{desc}} = a_c/2$

Initial Orbit (= Debris Orbit)

$$a_0 = a_t = R_{\oplus} + 1200 \text{ km}$$

$$e_0 = e_t = 0$$

$$i_0 = i_t$$

Final Orbit (= Parking Orbit)

$$a_{\text{fin}} = R_{\oplus} + 600 \text{ km}$$

$$e_{\text{fin}} = 0$$

$$i_{\text{fin}} = i_t$$

Descent Phase: blended control law

Objective

Find **optimal sail attitude** $\{\alpha_{opt}, \delta_{opt}\}$ that **minimizes** the **cost index** J at each time (MATLAB *fmincon*)

Cost Index

$$J = R_a \frac{d(a/a_0)}{dt} + R_e \frac{de}{dt} + R_i \frac{di}{dt}$$

$$R_a = \frac{a - a_{fin}}{|a_0 - a_{fin}|}, \quad R_e = e - e_{fin}, \quad R_i = i - i_{fin}$$

Weights

Necessary to redefine R_e and R_i since $e_0 = e_{fin} = 0$ and $i_0 = i_{fin}$

Table of Contents

- 1 Introduction
 - Mission Scenario
- 2 Dynamical Model
 - Equations of motion
 - Perturbing Accelerations
 - Eclipses
- 3 Transfer Strategy
 - 1st Phase: semimajor axis increase
 - 2nd Phase: blended control law
 - Descent Phase
- 4 Numerical Simulations
- 5 Conclusion

Numerical Simulations - 1st Phase - Data

Start Date: 01/01/2030

Characteristic **Acceleration**: $a_c = 0.1 \text{ mm/s}^2$

Stopping Criterion: $h_P = 1000 \text{ km}$

Sail parking orbit

$$h_0 = 600 \text{ km}$$

$$e_0 = 0$$

$$i_0 = 60 \text{ deg}$$

$$\Omega_0 = 0 \text{ deg}$$

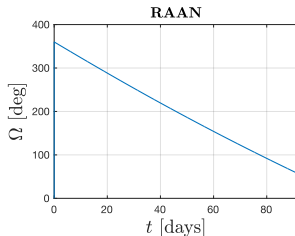
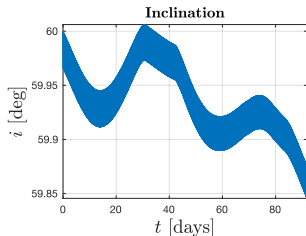
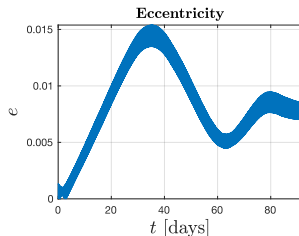
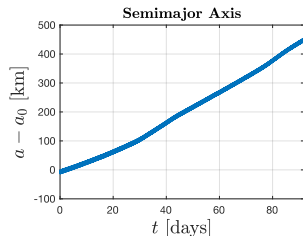
Target debris orbit

$$h_t = 1200 \text{ km}$$

$$e_t = 0$$

$$i_t = 60 \text{ deg}$$

$$\Omega_t = \Omega_{t,0} + \dot{\Omega}_t t$$

1st Phase - SMA increase - Plots

Results

$$\Delta t = 91.8 \text{ days}$$

$$\Delta a = 455.2 \text{ km}$$

$$\Delta e = 0.0074$$

$$\Delta i = -0.1251 \text{ deg}$$

Eclipse Time

$$\frac{\Delta t_{\text{ecl}}}{\Delta t} \approx 23.53 \%$$

2nd Phase - Debris targeting - Results

Genetic Algorithm Settings

Population Size = 50

Elite Count = 2

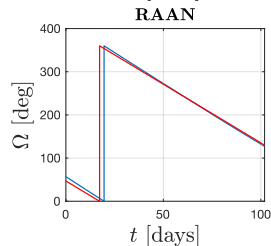
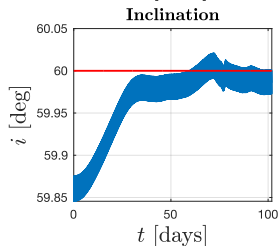
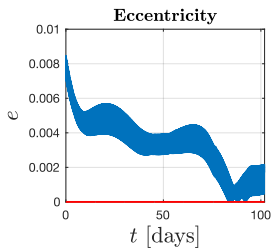
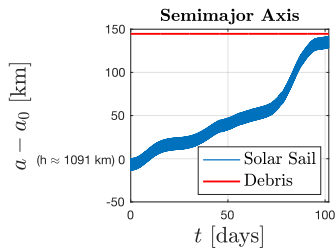
Number of Generations = 10

Function Tolerance = 1×10^{-6} $\Delta\Omega_0 = 10$ deg

GA Results

 $W_a = 0.27849822$ $W_e = 0.83082863$ $W_i = 0.76324954$ $F_{\text{opt}} = 1.0907 \times 10^{-6}$ **Flight Time** = 101.9 days

Orbit	a [km]	e	i [deg]	Ω [deg]
Sail initial	7433.5	0.0074	59.87	56.98
Debris initial	7578.1	0	60	46.98
Debris final	7578.1	0	60	130.07
Sail final Approx	7588.7	0.0009	59.99	130.05
Sail final Exact	7574.4	0.0017	60	127.23

2nd Phase - Debris targeting - Plots

Results

$$\Delta t = 101.9 \text{ days}$$

$$\Delta a = -3.7722 \text{ km}$$

$$\Delta e = 0.0017$$

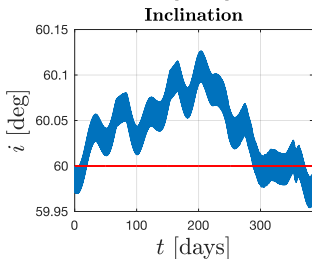
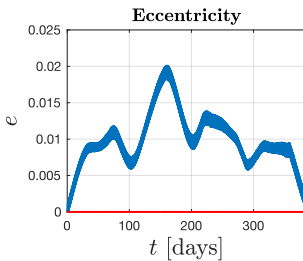
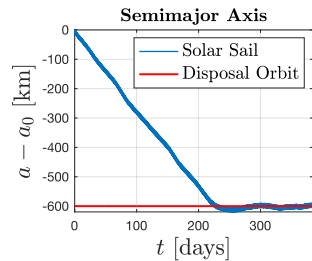
$$\Delta i = -0.0002 \text{ deg}$$

Right Ascension

$$\Delta \Omega_0 = 10 \text{ deg}$$

$$\Delta \Omega_{\text{fin}} \approx 2.84 \text{ deg}$$

Descent Phase - Plots



Results

 $\Delta t = 384.4$ days $\Delta a = -1.8550$ km $\Delta e = 0.0009$ $\Delta i = -0.0361$ deg

Table of Contents

- 1 Introduction
 - Mission Scenario
- 2 Dynamical Model
 - Equations of motion
 - Perturbing Accelerations
 - Eclipses
- 3 Transfer Strategy
 - 1st Phase: semimajor axis increase
 - 2nd Phase: blended control law
 - Descent Phase
- 4 Numerical Simulations
- 5 Conclusion

Conclusion

Conclusions

- **Locally-optimal laws work well** in perturbed environments
- **Blending seems effective** in controlling more orbital parameters at the same time
- **Approximate** model is pretty accurate for a **preliminary study**

Model Improvements

- Way to **estimate** the **RAAN** at the end of the 1st phase
- **Increase** number of **generations** in the Genetic Algorithm

Further Developments

Further Developments

- **Switching point** between 1st and 2nd phase to be **optimized**
- **Higher performance sails** to reduce parking orbit's height
→ **shorter re-entry time** for debris
- **Multiple debris** removal strategy
→ best use of solar sails as **propellantless** devices
→ reduction of **launch cost**

Thank you for your attention!

Christian Bianchi

christian.bianchi@phd.unipi.it

Department of Civil and Industrial Engineering

University of Pisa, Pisa (PI), Italy